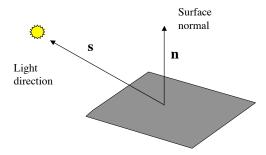
Lambertian Reflection



Brightness is proportional to nos

Lambertian Reflection

Why is brightness proportional to **n•s**?

Intuitive argument: The surface scatters light in all directions equally, but as the angle of the light becomes oblique, the amount of light per unit area is reduced (foreshortening) by a factor of the cosine of the angle.



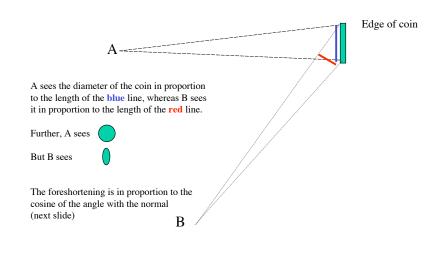
Comments on light source direction

The direction to a nearby light changes as you move around in the scene.

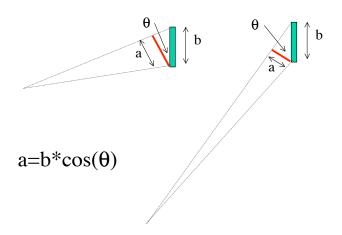
If we say a light source is "at infinity", we mean that it is so far away that only the direction is important.

Example: On the scale of a city, the sun is at infinity.

Foreshortening illustrated

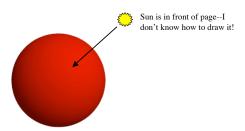


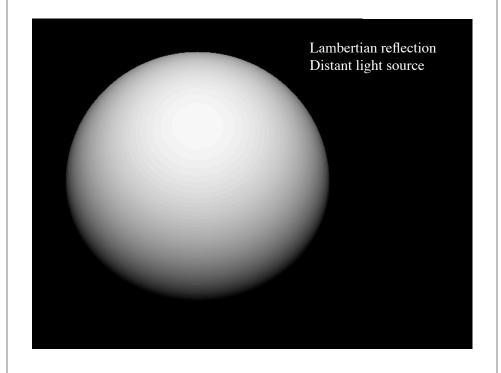
Foreshortening illustrated



Lambertian surfaces

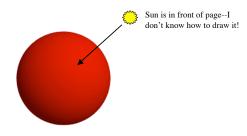
• Surface brightness is only a function of the foreshortening of the incident light (the more oblique it is, the less bright the surface).



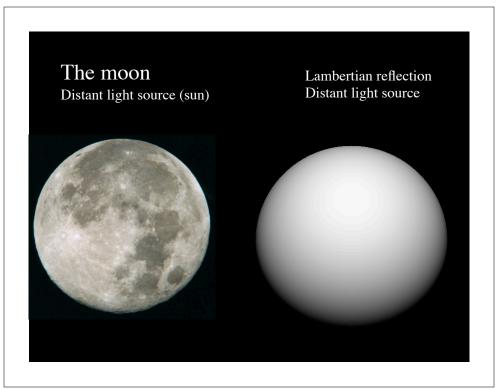


Lambertian surfaces

• Surface brightness is only a function of the foreshortening of the incident light (the more oblique it is, the less bright the surface).



• Question: Is the moon a Lambertian reflector?



Lambertian Reflection

Most the world is not Lambertian

Lambertian assumption failures

Lambertian Reflection

Most the world is not Lambertian

Lambertian assumption failures

Rough surfaces--important example--the moon is not Lambertian

Dielectrics (plastics, many paints)

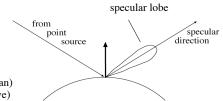
Metallic surfaces

Skin

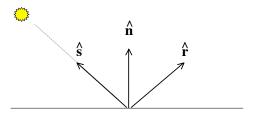
Specular surfaces

- Another important class of surfaces is specular (mirror-like).
 - specular surfaces reflect a significant amount of energy in the specular (mirror) direction
 - produces "highlights"
- · Two related cases
 - a perfect mirror
 - a fuzzy mirror

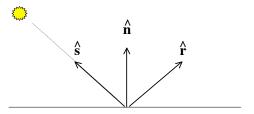
• Typically there is a diffuse (Lambertian) component as well (effects are additive)



Computing reflection (specular) direction



Computing reflection (specular) direction



$$\hat{\mathbf{s}} + \hat{\mathbf{r}} = k\hat{\mathbf{n}}$$

and

$$\hat{\mathbf{n}} \cdot \hat{\mathbf{s}} = \hat{\mathbf{n}} \cdot \hat{\mathbf{r}}$$

$$\hat{\mathbf{n}} \cdot \hat{\mathbf{s}} + \hat{\mathbf{n}} \cdot \hat{\mathbf{r}} = k \implies k = 2\hat{\mathbf{n}} \cdot \hat{\mathbf{s}}$$

So
$$\hat{\mathbf{r}} = 2(\hat{\mathbf{n}} \cdot \hat{\mathbf{s}})\hat{\mathbf{n}} - \hat{\mathbf{s}}$$

Phong's model of specularities

- There are very few cases where the exact shape of the specular lobe matters.
- Typically:
 - very, very small --- mirror
 - small -- blurry mirror
 - bigger -- see only light sources as "specularities"
 - very big -- faint specularities



- reflected energy falls off with

$$\cos^n(\delta\vartheta)$$







Plus Specular Highlight

from

δΘ

direction

specular

http://www.geocities.com/SiliconValley/Horizon/6933/shading.html