

Homogenous linear least squares

Example 3.1 in book

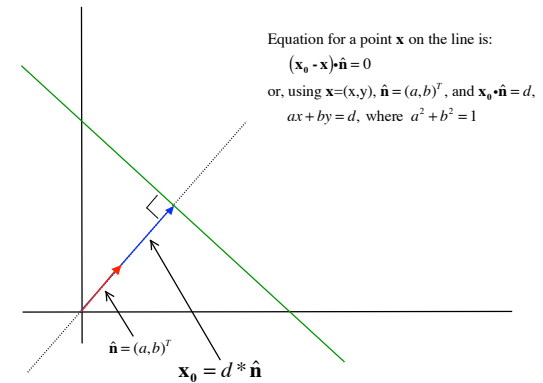
(fitting a line to points, a better way for many applications)

Key initial point: The perpendicular distance from a point x_i , to a line $ax+by=d$, where $a^2+b^2=1$ is given by:

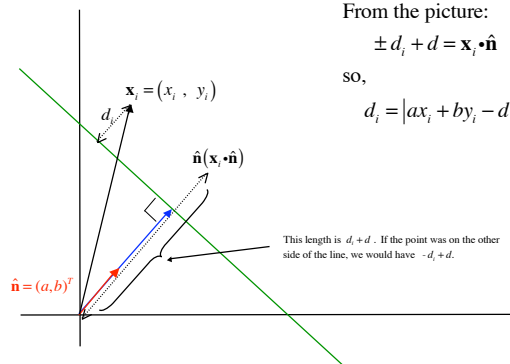
$$d_i = |ax_i + by_i - d|$$

(See next two slides for geometry)

Line Fitting



Line Fitting



Line Fitting (continued)

$$E = \sum d_i^2 = \sum (d - ax_i - by_i)^2$$

$$\frac{\partial E}{\partial d} = 2 \sum (d - ax_i - by_i) = 0$$

$$\text{So, } d = a\bar{x} + b\bar{y}$$

Line Fitting (continued)

$$d = a\bar{x} + b\bar{y} \quad (\text{from previous slide})$$

$$\begin{aligned} E &= \sum (a\bar{x} - ax_i + b\bar{y} - by_i)^2 \\ &= \sum ((\bar{x} - x_i, \bar{y} - y_i) \bullet (a, b))^2 \\ &= |U\mathbf{n}|^2, \text{ where } U = \begin{pmatrix} \bar{x} - x_1 & \bar{y} - y_1 \\ \dots & \dots \\ \bar{x} - x_n & \bar{y} - y_n \end{pmatrix} \end{aligned}$$

So, we solve $U\mathbf{n}=0$ in the least squares sense, with $a^2 + b^2 = 1$

Back to cameras (§3.2.1)

Goal one: Find the matrix M linking world coordinates to image coordinates from image of calibration object.

$$\begin{pmatrix} U \\ V \\ W \end{pmatrix} = M \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix} = MP$$

Recall, that since the above is in terms of homogeneous coordinates we have to work in terms of the observed image coordinates, $u=U/W$ and $v=V/W$

Recall that we form column vectors from the rows of M and stack the columns on top of one another to get the vector of unknowns, \mathbf{m} .

Recall that we derived the following equation for \mathbf{m} , to be solved subject to $|\mathbf{m}|=1$ in the least squares sense.

$$\begin{pmatrix} \mathbf{P}_1^T & -u_1\mathbf{P}_1^T \\ & \mathbf{P}_1^T & -v_1\mathbf{P}_1^T \\ \dots & \dots & \dots \\ \mathbf{P}_i^T & -u_i\mathbf{P}_i^T \\ & \mathbf{P}_i^T & -v_i\mathbf{P}_i^T \\ \dots & \dots & \dots \\ \mathbf{P}_n^T & -u_n\mathbf{P}_n^T \\ & \mathbf{P}_n^T & -v_n\mathbf{P}_n^T \end{pmatrix} \begin{pmatrix} \mathbf{m}_1^T \\ \mathbf{m}_2^T \\ \mathbf{m}_3^T \end{pmatrix} = \mathbf{0}$$

So, now we can simply apply the eigenvalue method in the previous slides to solve for \mathbf{m} .

Intrinsic/extrinsic parameters

Recall goal two: Given M , recover the intrinsic parameters.

See §3.2.2 for the development of some formulas. Grad students will use a simplified version of them in assignment three (relatively straight forward, but a bit complex)

Lambertian surfaces

Simple rule for shading--attenuate brightness by

$$\mathbf{n} \cdot \mathbf{s}$$

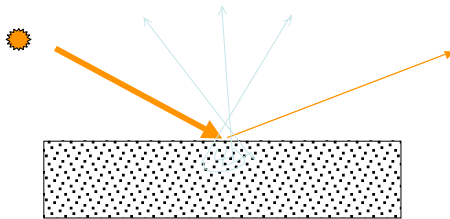
Surface
normal

Light source
direction

Must know this

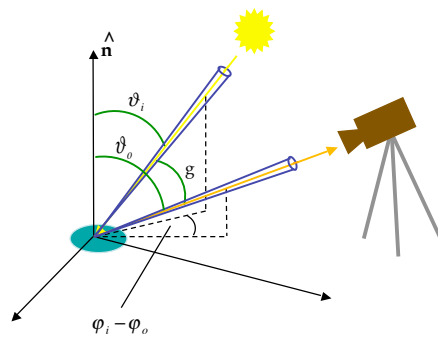
Light interacting with the world

- The light captured by camera carries information about what is in the world **because** what is in the world interacts with it differently depending on what it is.
- Many effects when light strikes a surface. It could be:
 - absorbed
 - transmitted
 - reflected
 - scattered (in a variety of directions!)



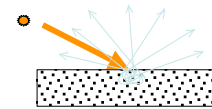
Bidirectional Reflectance Distribution Function (BRDF)

- The BRDF is a technical way of specifying how light from sources interacts with the matter in the world
- Understanding images requires understanding that this varies as a function of materials. The following “look” different
 - mirrors
 - white styrofoam
 - colored construction paper
 - colored plastic
 - gold
- The BRDF is the **ratio** of what comes out to what came in
- What comes out <--> “radiance”
- What goes in <--> “irradiance”
- Details on the BRDF available as supplementary material



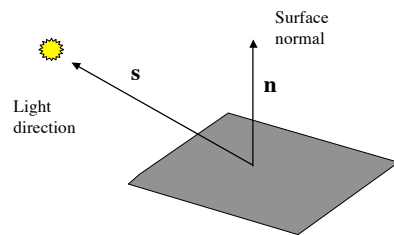
Lambertian surfaces

- Simple special case of reflectance: ideal diffuse or matte surface--e.g. cotton cloth, matte paper.
- Surface appearance is independent of viewing angle.
- Typically such a surface is the result of lots of scattering---the light "forgets" where it came from, and it could end up going in any random direction.



- What counts is how much light power reaches the surface

Lambertian Reflection

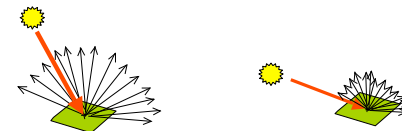


Brightness is proportional to $\mathbf{n} \cdot \mathbf{s}$

Lambertian Reflection

Why is brightness proportional to $\mathbf{n} \cdot \mathbf{s}$?

Intuitive argument: The surface scatters light in all directions equally, but as the angle of the light becomes oblique, the amount of light per unit area received is reduced (foreshortened) by a factor of the cosine of the angle.



The same light is spread over a , giving intensity, i_a , as is spread over b , giving intensity, i_b . This means that:

$$a \cdot i_a = b \cdot i_b$$

or, because a is the length of the perpendicular,

$$i_b = i_a \left(\frac{a}{b} \right) = i_a \cos(\theta)$$



The sun is far away so light rays are nearly parallel.

